

$$V_{rms}^{H_2} = \sqrt{\frac{3 \cdot 8.31 \cdot 293}{2 \cdot 10^{-3}}} = 1910 \text{ m/s}$$

If earth a small asteroid - gas escapes

(skip intermolecular collisions)

Heat Capacity (constant volume).

Monatomic Ideal Gas: all stored energy is translational kinetic energy!

total energy $\rightarrow E = K = \frac{3}{2} nRT = \frac{3}{2} NkT$

$$dQ = n C_v dT = dE = \frac{3}{2} n R dT$$

constant volume

$$C_v = \frac{3}{2} R$$

per mole

ideal monatomic gas.

Diatomic Molecule

$\frac{1}{2}kT$ $\frac{1}{2}kT$ $\frac{1}{2}kT$
 transl x y z

$\frac{1}{2}kT$ $\frac{1}{2}kT$
 2 degrees of freedom, ROTATION!
 3rd axis - unmeasurable.

$$E_1 = \frac{5}{2} kT$$

per mole: $E = N_A E_1 = \frac{5}{2} (N_A \cdot k) T$

$$E = \frac{5}{2} RT$$

diatomic
 $C_v = \frac{5}{2} R$

Vibration

Quantum Mechanics
 "Freezes Out" these.

$\frac{1}{2}kT \ll$ Energy needed (Room Temp)

Solid: 3 translational

3 vibrational (solid generally not frozen out)

Fig 18.16

$$C_v \approx 3R$$

per mole, $\sim 25 \frac{J}{mol \cdot K}$

Molecular Speeds

- molecules have a variety of velocities
- What is the distribution of...
velocities, speeds

↑
3-d
+ - !

↑
 ≥ 0 , 1 dimensional
Test scores

- Boltzmann Factor... a molecule
→ in thermal equilibrium.
→ temperature T (absolute)

1 molecule · $P(E) \propto e^{\frac{-E}{kT}}$

monatomic gas: $E = \frac{1}{2}m(v_x^2 + v_y^2 + v_z^2)$

$$P(v_x, v_y, v_z) \propto e^{\frac{-m(v_x^2 + v_y^2 + v_z^2)}{2kT}}$$

$$M = N_A k \rightarrow e^{\frac{-M(v_x^2 + v_y^2 + v_z^2)}{2RT}}$$

$$\propto e^{\dots}$$

$$\dots$$