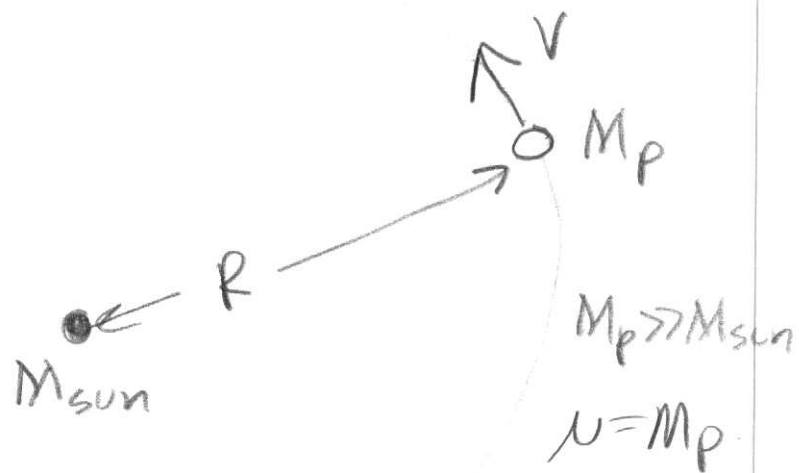


Slingshot Effect

First ... look at circular motion
and ponder it....

Planet about the sun..

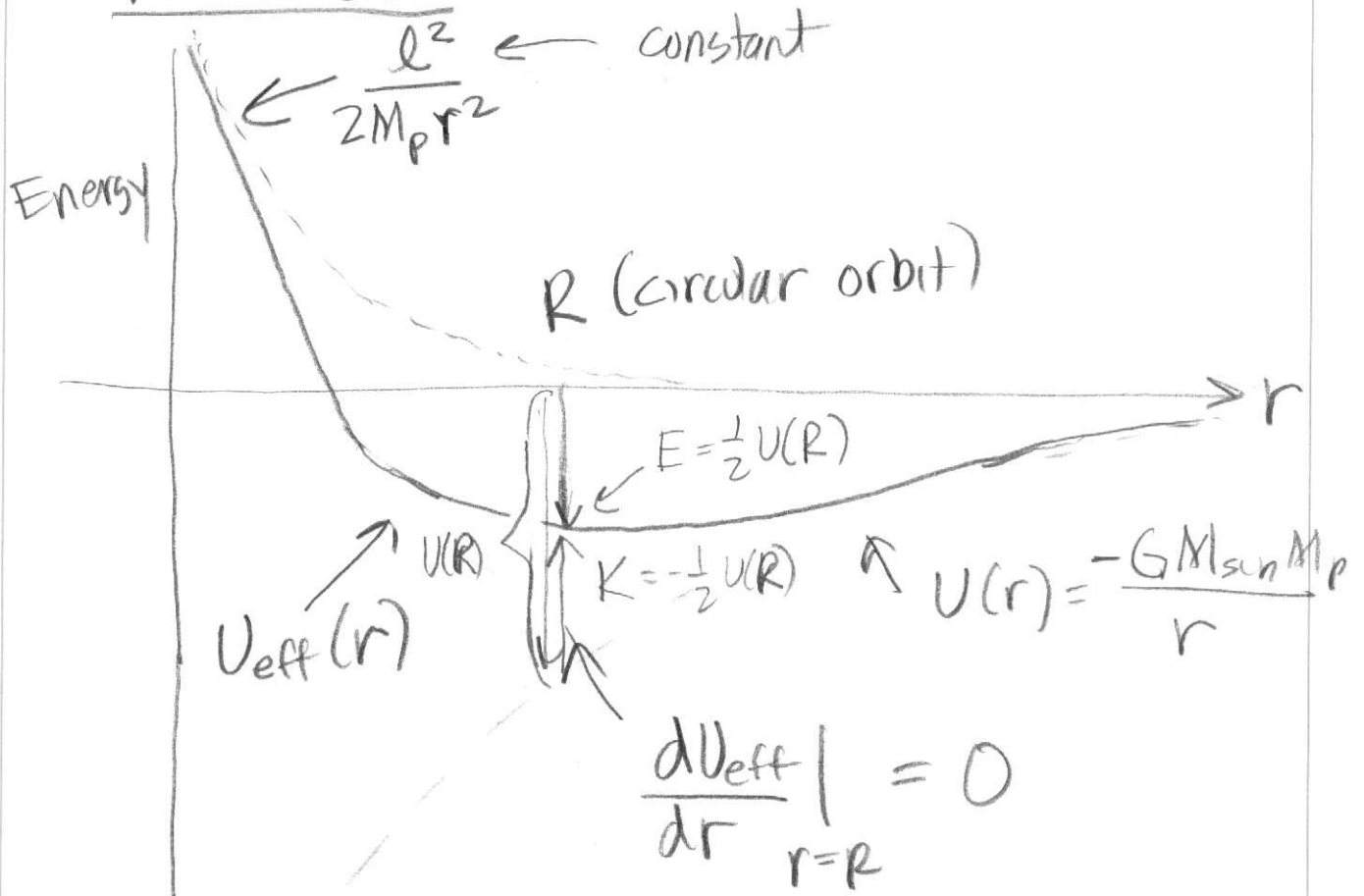


Simply : $-M_P \cdot \frac{v^2}{R} = -G \frac{M_{\text{sun}} M_P}{R^2}$

so, $K = \frac{1}{2} M_P v^2 = \frac{1}{2} \cdot \underbrace{G \frac{M_{\text{sun}} M_P}{R}}_{-U(R)}$

$$K = \frac{1}{2} \cdot (-U(R))$$

Visualize:



$$-\frac{l^2}{M_p R^3} + \frac{GM_{\text{sun}} M_p}{R^2} = 0$$

$$\frac{(M_p R v)^2}{M_p R^3} = \frac{GM_{\text{sun}} M_p}{R^2}$$

same
as
before! $\rightarrow$

$$\frac{1}{2} M_p v^2 = \frac{1}{2} \frac{GM_{\text{sun}} M_p}{R}$$

Second

How to use a planet to escape the sun....



Sun

tiny spaceship
at rest!



Planet
(trapped)

Collision
initial

Planet
 M_p



v

spaceship
 $M_p \gg M_{\text{spaceship}}$


final

v $2v$



(Best Case!)

↑
spaceship...
is it bound?

$$K_{\text{spaceship}} = \frac{1}{2} M_{\text{spaceship}} (2v)^2 = 4 \cdot \frac{1}{2} M_s v^2$$

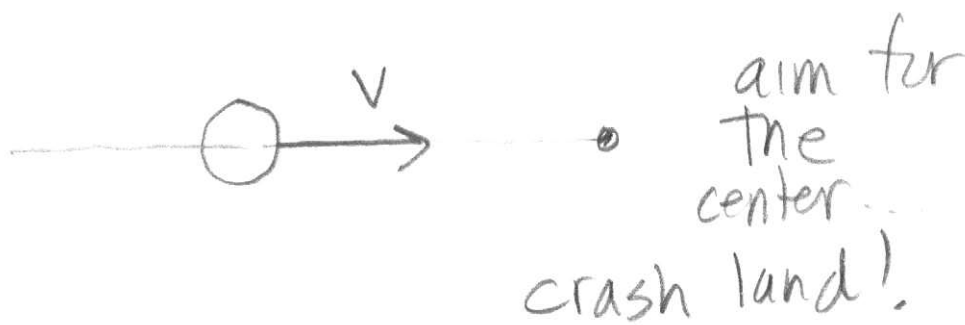
$$K_s = 4 \times \frac{1}{2} M_s v^2 = 4 \times \left(-\frac{1}{2} U(R) \right)$$

$$E = K_s + U(R) = -2U(R) + U(R) = -U(R)$$

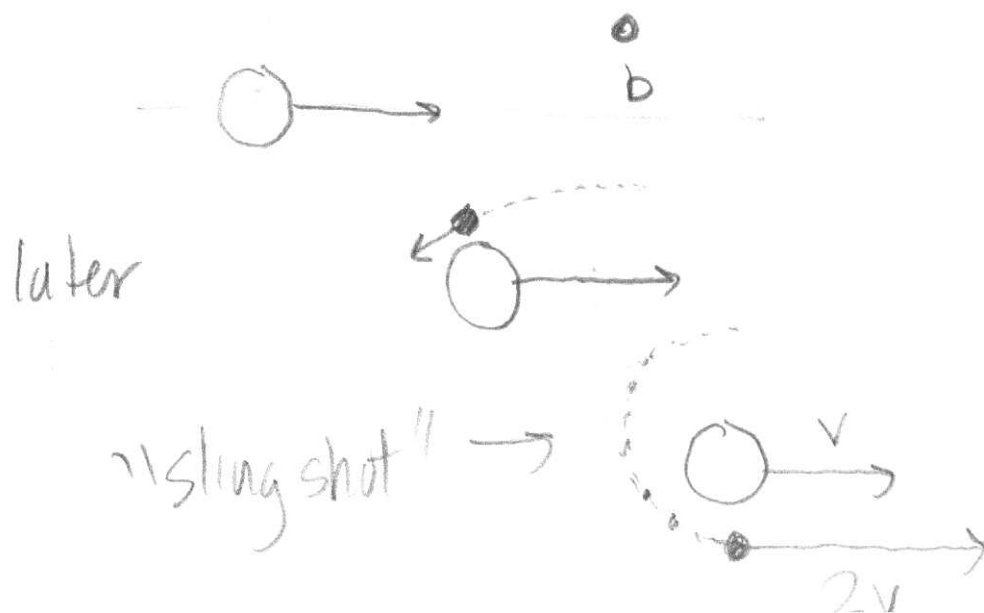
$$E = - \left(-\frac{GM_s M_{\text{sun}}}{R} \right) = + \frac{GM_s M_{\text{sun}}}{R} > 0$$

$\Rightarrow$ spaceship can escape the sun!

HOW:



Want to plunk yourself down off center a bit..



Full Solution

$$E = \frac{1}{2} \mu \dot{r}^2 + \frac{l^2}{2\mu r^2} + U(r) = \text{constant}$$

$$\dot{r} = \frac{dr}{dt} = \pm \left[\frac{2}{\mu} \left(E - \frac{l^2}{2\mu r^2} - U(r) \right) \right]^{1/2}$$

$$l = m r^2 \dot{\theta} \Rightarrow \dot{\theta} = \frac{d\theta}{dt} = \frac{l}{\mu r^2} \leftarrow \text{constant}$$

$$\frac{\frac{d\theta}{dt}}{\frac{dr}{dt}} = \frac{d\theta}{dr} = \frac{\frac{l}{\mu r^2}}{\pm \left[\frac{2}{\mu} \left(E - \frac{l^2}{2\mu r^2} - U(r) \right) \right]^{1/2}}$$

$$\frac{d\theta}{dr} = \pm r \frac{l}{\left[\mu^2 r^2 \cdot \frac{2}{\mu} \left(E - \frac{l^2}{2\mu r^2} - U(r) \right) \right]^{1/2}}$$

$$= \pm \frac{l}{r \left[2\mu r^2 \left(E - \frac{l^2}{2\mu r^2} - U(r) \right) \right]^{1/2}}$$

$$\frac{d\theta}{dr} = \pm \frac{l}{r \left[(2\mu E) r^2 + (2\mu C) r - l^2 \right]^{1/2}}$$

$\pm$ / clockwise / counterclockwise.