

$$\left(\frac{\hbar^2}{2\mu} \left[-\frac{d^2}{dr^2} + \frac{l(l+1)}{r^2} \right] + V(r) \right) U_{El}(r) = E U_{El}(r)$$

now, diverge a little from the book (really, just take things in a slightly different order) GENERAL PROPERTIES: 1) l

- $l=0$.. no orbital angular momentum

Almost just like the 1-d Schrödinger:

$$\left(\frac{-\hbar^2}{2\mu} \frac{d^2}{dr^2} + V(r) \right) U_{E0}(r) = E U_{E0}(r)$$

★ Simplify further: $V(r)=0$!

$$\frac{-\hbar^2}{2\mu} \frac{d^2}{dr^2} U_{E0}^f(r) = E U_{E0}^f(r) \quad f = \text{"free"}$$

with $k^2 \equiv \frac{2\mu E}{\hbar^2}$ it would seem that $U_{E0}^f(r) = A e^{ikr} + B e^{-ikr}$

$A + B \sim$ arbitrary complex numbers

is a solution.

★ Big point: radial 3-d geometry that includes the origin ($r=0$) eliminates all but $A = -B$

Why? An odd singularity creeps in; one has to go back to $R_{E0}^f(r)$ to see it:

$$R_{E0}^f(r) = \frac{U_{E0}^f}{r} = \frac{(A e^{ikr} + B e^{-ikr})}{r}$$

Since $\Psi_{E0}^f(\vec{x}) = R_{E0}^f(r) Y_0^0(\theta, \phi) = \frac{1}{\sqrt{4\pi}} R_{E0}^f(r)$
 is just a multiplicative constant different,
 we can go back to the full SE:

$$-\frac{\hbar^2}{2m} \nabla^2 R_{E0}^f(r) = E R_{E0}^f(r)$$

$$-\frac{\hbar^2}{2m} \nabla^2 \left[\frac{Ae^{ikr} + Be^{-ikr}}{r} \right] = E \left[\frac{Ae^{ikr} + Be^{-ikr}}{r} \right]$$

if one checks both sides of this equation, it almost works; the point $r=0$ has a problem on the left hand side, however:

$$\nabla^2 \left(\frac{1}{r} \right) = -4\pi \delta^{(3)}(\vec{x}) \quad \left\{ \begin{array}{l} \text{from} \\ \text{Electromagnetism} \end{array} \right.$$

the right hand side has nothing like a term such as this. The remedy is to make the coefficient of the $\delta^{(3)}(\vec{x})$ zero near $r=0$, or

$$A+B=0, \quad B=-A$$

★ Conclusion: $U_{E0}^f(r) = A \sin(kr)$
 (if $r=0$ present)

• $l \neq 0 \dots + \frac{l(l+1)}{r^2}$ is a barrier near the origin.. particles pushed out.

★ analyze case where $V(r)$ is negligible compared to $l(l+1)/r^2$ as $r \rightarrow 0$

E also becomes unimportant

$$\left[-\frac{d^2}{dr^2} + \frac{l(l+1)}{r^2} \right] U_l(r) = 0$$

Try: $U_\ell(r) \propto r^\alpha$ S.E. gives:

$$-\alpha(\alpha-1) + \ell(\ell+1) = 0$$

$$\alpha^2 - \alpha - \ell(\ell+1) = 0$$

$$\alpha = \frac{1 \pm \sqrt{1 + 4\ell(\ell+1)}}{2} = \frac{1 \pm (2\ell+1)}{2}$$

$$\alpha = \ell+1, -\ell$$

very important

\rightarrow $U_\ell(r) \propto r^{\ell+1}$ near origin
suppression as expected Type-A

$U_\ell(r) \propto \frac{1}{r^\ell}$ enhancement near origin Type-B ($\ell \geq 1$)

When $r=0$ is in the problem, we keep only Type-A solutions.

Looking back at $\ell=0$, we see that $U_\ell(r) \propto r^1$, the Type-A solution, was in fact selected by the argument we made.

For $\ell \geq 1$,

The enhancement at the origin is actually (physically) sufficient to eliminate Type-B. An additional flaw, that Type-B solutions result in a non-Hermitian Hamiltonian, is detailed on p. 341 of text.

Note: if $r=0$ is not to be considered, Type-II solutions are fine!

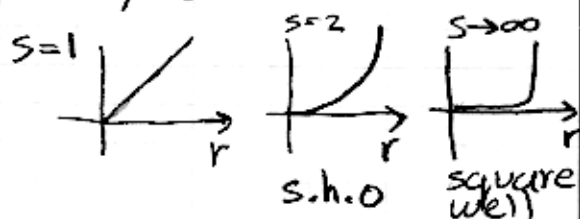
GENERAL PROPERTIES: 2) Length Scale, Power Law Potentials

Power Law Potential:

$$V(r) = V_0 r^s \quad \bullet \text{ attractive: a) } V_0 > 0 \quad s > 0$$

$s = \text{power}$

- dimensions of V_0 are $\frac{\text{Energy}}{\text{Length}^s}$



b) $V_0 < 0 \quad s < 0$



• repulsive! reverse V_0 sign

$$\left\{ \frac{\hbar^2}{2\mu} \left[-\frac{d^2}{dr^2} + \frac{l(l+1)}{r^2} \right] + V_0 r^s \right\} U_{El}(r) = E U_{El}(r)$$

define a length scale a , so a dimensionless radius $\rho \equiv \frac{r}{a}$; $r = a\rho$

plugging back in:

$$\left\{ \frac{\hbar^2}{2\mu a^2} \left[-\frac{d^2}{d\rho^2} + \frac{l(l+1)}{\rho^2} \right] + V_0 a^s \rho^s \right\} U_{El}(\rho) = E U_{El}(\rho)$$

↑ Idea: choose a so that these coefficients are equal.

Physics: a is where radial contribution to kinetic energy = potential energy.

$$\frac{\hbar^2}{2\mu a^2} = |V_0| a^s \Rightarrow$$

$$a = \left(\frac{\hbar^2}{2\mu |V_0|} \right)^{\frac{1}{s+2}}$$

- $s = 2$
 $|V_0| = \frac{1}{2} \mu \omega^2$

$$a = \left(\frac{\hbar^2}{2\mu \cdot \frac{1}{2} \mu \omega^2} \right)^{\frac{1}{4}} = \sqrt{\frac{\hbar}{m\omega}}$$

see p. 191, 7.3.6

- $s = -1$
 $|V_0| = Ze^2$
 (charge Z)

$$a = \left(\frac{\hbar^2}{2\mu Ze^2} \right)^{-\frac{1}{-1+2}}$$

$$= \frac{1}{2Z} \frac{\hbar^2}{\mu e^2} = \frac{1}{2Z} a_0$$

$a_0 \rightarrow$ Bohr radius
 see p. 357, 13.1.24

- $s \rightarrow \infty$

$$a \rightarrow \left(\frac{\hbar^2}{2\mu |V_0|} \right)^{\frac{1}{\infty+2}} \rightarrow 0$$

but $|V_0|$ can be simultaneously adjusted to make a anything

What is the value of the common coefficients?

$$\frac{\hbar^2}{2\mu a^2} = |V_0| a^s = |V_0| \left(\frac{\hbar^2}{2\mu |V_0|} \right)^{\frac{s}{s+2}} \equiv E_0$$

in these terms, the radial equation now becomes:

$$\left\{ -\frac{d^2}{dp^2} + \frac{l(l+1)}{p^2} + \text{sgn}(V_0) p^s \right\} U_{\ell\ell}(p) = \epsilon U_{\ell\ell}(p)$$

"dimensionless radial equation"
 \uparrow
 $+1$ when $V_0 > 0$
 -1 when $V_0 < 0$